# Scalar Waves on a Naked Singularity Background 

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Class. Quant. Grav. 2004

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- There are no troublesome indices.
- People (mostly Wald ${ }^{2}$ and students) suggest well-posedness of wave equations as a substitute for geodesic completeness.

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More generally,

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- Dispersive Estimates ( $L^{\infty}$ decay):

$$
\|u(t)\|_{L^{\infty}\left(\mathbf{R}^{3}\right)} \leq C t^{-1}\left(\|\nabla u(0)\|_{L^{1}\left(\mathbf{R}^{3}\right)}+\left\|\partial_{t} u(0)\right\|_{L^{1}\left(\mathbf{R}^{3}\right)}\right)
$$

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One nice thing about this Strichartz is that it is Lorentz invariant. Strichartz estimates, like dispersive estimates, can be used to prove stability for non-linear wave equations, but Strichartz estimates are often true in contexts where dispersive estimates fail.

## Scalar Wave Equation in Spherical Symmetry

Metric in isothermal coordinates:

$$
g_{\mu \nu} d x^{\mu} d x^{\nu}=\alpha(r)^{2}\left(-d t^{2}+d r^{2}\right)+\rho(r)^{2}\left(d \varphi^{2}+\sin ^{2} \varphi d \theta^{2}\right)
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(Massless, Chargeless) Scalar Wave Equation:

$$
g^{\mu \nu} \psi_{; \mu \nu}=0
$$

or

$$
\partial_{t}^{2} \psi-\frac{1}{\rho^{2}} \partial_{r}\left(\rho^{2} \partial_{r} \psi\right)-\frac{\alpha^{2}}{\rho^{2}} \Delta_{\mathrm{Sph}} \psi=0 .
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u=\rho \psi / r
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The coefficient in front of $\Delta_{\mathrm{Sph}} u$ is $\alpha^{2} / \rho^{2}$ rather than $1 / r^{2}$, but we can ignore that for spherically symmetric solutions. Compared to the scalar wave equation in Minkowski space, there is an additional scalar potential

$$
V(r)=\rho^{\prime \prime}(r) / \rho(r)
$$

## Reissner-Nordström

For Reissner-Nordström,

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\rho^{\prime}(r)=\alpha^{2} \quad \alpha=\sqrt{1-\frac{2 m}{\rho}+\frac{e^{2}}{\rho^{2}}}
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in natural units, $G=c=1$.

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- If $|e|=m$, the extremal case, then the quadratic has a double root at $\rho=m$, which is again a horizon and the metric is valid outside the horizon.
- If $|e|>m$, the super-extremal case, then there are no horizons and the metric above is valid for all $r>0$, but is highly singular at $r=0$.


## Super-extremal Case

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A u=-\frac{1}{r^{2}} \partial_{r}\left(r^{2} \partial_{r} u\right)-\frac{\alpha^{2}}{\rho^{2}} \Delta_{\mathrm{Sph}} u+V(r) u
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but with what domain?

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Almost all symmetric differential operators appearing in Mathematical Physics are essentially self-adjoint.

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An alternate characterisation, due to Krein, is this: If $A_{F}$ is the Friedrichs extension and $A_{E}$ is any other positive self-adjoint extension of $A$ then

$$
u \in \operatorname{Dom}\left(A_{F}\right) \Longrightarrow u \in \operatorname{Dom}\left(A_{E}\right) \text { and }\left\langle u, A_{F} u\right\rangle \leq\left\langle u, A_{E} u\right\rangle
$$

[^7]
## BPST-Z

Now that we have a well defined problem, we can use a theorem we proved earlier with Burq and Planchon: ${ }^{4}$
Let $V \in C^{1}\left(\mathbf{R}^{+}\right)$satisfy

- $\sup _{r \in \mathbf{R}^{+}} r^{2} V(r)<\infty$
- $\inf _{r \in \mathbf{R}^{+}} r^{2} V(r)>-1 / 4$
- $\sup _{r \in \mathbf{R}^{+}} r^{2} \frac{d}{d r}(r V(r))<1 / 4$,
let $P=-\Delta+V$, and let $P_{F}$ be the Friedrichs extension of $P$. Then there exists a $C$ such that if

$$
\partial_{t}^{2} u+P_{F} u=0
$$

then

$$
\|u\|_{L^{4}\left(\mathbf{R}^{1+3}\right)} \leq C \mathbf{E}_{1 / 2}[u]
$$

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## What needs to be done?

There are two things to be checked, before we can apply the theorem:

- We want a theorem with natural (in terms of the R-N metric) norms on $\psi$, not the usual (Minkowski) norms on $u$.


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and $\frac{\alpha r}{\rho}$ is bounded.

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- We need to check the three hypotheses on our $V$.


## Checking V

Recall:

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We can eliminate one variable by scaling.

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r=\rho-r_{0}+m \log \left(\frac{\rho^{2}-2 m \rho+e^{2}}{e^{2}-m^{2}}\right)+\frac{2 m^{2}-e^{2}}{\sqrt{e^{2}-m^{2}}} \arctan \frac{\rho-m}{\sqrt{e^{2}-m^{2}}}, \\
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We can eliminate one variable by scaling.
We still have nasty transcendental functions of the two variables $r / e$ and $m / e$ whose zeroes we need to find.

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Doesn't quite work as written, but something very similar does, at least for $|e| \geq 2 m$.

## Algebra

Eventually, after much algebra, the problem reduces to one of real algebraic geometry:
Is the curve

$$
\begin{aligned}
& 72 x^{14}-\left(576+432 y^{2}\right) x^{13}+\left(1947+3552 y^{2}+1152 y^{4}\right) x^{12} \\
& -\left(3504+11988 y^{2}+10464 y^{4}+1440 y^{6}\right) x^{11} \\
& +\left(3452+20360 y^{2}+38762 y^{4}+15384 y^{6}+720 y^{8}\right) x^{10} \\
& -\left(1536+16456 y^{2}+71800 y^{4}+66316 y^{6}+10536 y^{8}\right) x^{9} \\
& +\left(2040 y^{2}+62966 y^{4}+143492 y^{6}+57803 y^{8}+2160 y^{10}\right) x^{8} \\
& -\left(-4608 y^{2}+8608 y^{4}+153832 y^{6}+154672 y^{8}+21648 y^{10}\right) x^{7} \\
& +\left(-20100 y^{4}+48272 y^{6}+208760 y^{8}+83120 y^{10}+2760 y^{12}\right) x^{6} \\
& -\left(-36120 y^{6}+104440 y^{8}+151552 y^{10}+20824 y^{12}\right) x^{5} \\
& +\left(-33769 y^{8}+109100 y^{10}+58958 y^{12}+1908 y^{14}\right) x^{4} \\
& -\left(-17900 y^{10}+62848 y^{12}+11680 y^{14}\right) x^{3} \\
& +\left(-5530 y^{12}+20912 y^{14}+944 y^{16}\right) x^{2} \\
& -\left(-972 y^{14}+3888 y^{16}\right) x+\left(-81 y^{16}+324 y^{18}\right)=0
\end{aligned}
$$

## Algebra (Continued)

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Definition: The edge polynomial associated with an (oriented) edge of the Newton polygon is the polynomial in one variable whose coefficients are the numbers associated to the lattice points in that edge, taken in order.

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Note that there is a classical algorithm for checking the existence of real zeroes of polynomials in one variable, the Sturm test.

[^11]
## So now what?

For the scalar wave equation in super-extremal Reissner-Nordström,

- energy estimates are trivial, once you figure out how to get a well-defined problem,
- Strichartz estimates, somewhat miraculously, are true, at least for spherically symmetric data,
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There are some interesting new non-linear stability results for MBI, by Speck.

Questions?


[^0]:    ${ }^{2}$ J. Math. Phys. 1980

[^1]:    ${ }^{2}$ J. Math. Phys. 1980

[^2]:    ${ }^{2}$ J. Math. Phys. 1980

[^3]:    ${ }^{2}$ J. Math. Phys. 1980

[^4]:    ${ }^{2}$ J. Math. Phys. 1980

[^5]:    ${ }^{2}$ J. Math. Phys. 1980

[^6]:    ${ }^{3}$ See Seggev, 2003.

[^7]:    ${ }^{3}$ See Seggev, 2003.

[^8]:    ${ }^{4}$ Indiana J of Math, 2004, but use Arxiv instead!

[^9]:    ${ }^{5}$ Forum Mathematicum, 2007

[^10]:    ${ }^{5}$ Forum Mathematicum, 2007

[^11]:    ${ }^{5}$ Forum Mathematicum, 2007

